### 6. Ray Tracing

### Reading

#### Required:

- Watt, sections 1.3-1.4, 12.1-12.5.1.
- T. Whitted. An improved illumination model for shaded display. Communications of the ACM 23(6), 343-349, 1980. [In the reader.]

#### Further reading:

- A. Glassner. An Introduction to Ray Tracing. Academic Press, 1989. [In the lab.]
- K. Turkowski, "Properties of Surface Normal Transformations," Graphics Gems, 1990, pp. 539-547. [In the reader.]

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## **Geometric optics**

Modern theories of light treat it as both a wave and a particle.

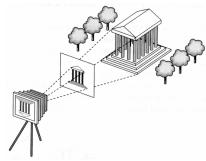
We will take a combined and somewhat simpler view of light – the view of **geometric optics**.

Here are the rules of geometric optics:

- Light is a flow of photons with wavelengths.
   We'll call these flows "light rays."
- Light rays travel in straight lines in free space.
- Light rays do not interfere with each other as they cross.
- Light rays obey the laws of reflection and refraction.
- Light rays travel form the light sources to the eye, but the physics is invariant under path reversal (reciprocity).

## Synthetic pinhole camera

The most common imaging model in graphics is the synthetic pinhole camera: light rays are collected through an infinitesimally small hole and recorded on an **image plane**.



For convenience, the image plane is usually placed in front of the camera, giving a non-inverted

Viewing rays emanate from the **center of projection** (COP) at the center of the lens (or pinhole).

The image of an object point *P* is at the intersection of the viewing ray through *P* and the image plane.

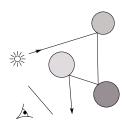
of the viewing ray through *P* and the image plane.

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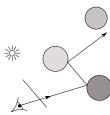
## Eye vs. light ray tracing

Where does light begin?

At the light: light ray tracing (a.k.a., forward ray tracing or photon tracing)



At the eye: eye ray tracing (a.k.a., backward ray tracing)

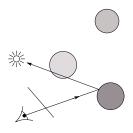


We will generally follow rays from the eye into the scene.

## **Precursors to ray tracing**

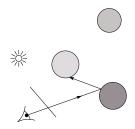
Local illumination

• Cast one eye ray, then shade according to light



Appel (1968)

• Cast one eye ray + one ray to light

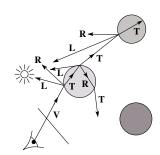


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# Whitted ray-tracing algorithm

In 1980, Turner Whitted introduced ray tracing to the graphics community.

- Combines eye ray tracing + rays to light
- Recursively traces rays

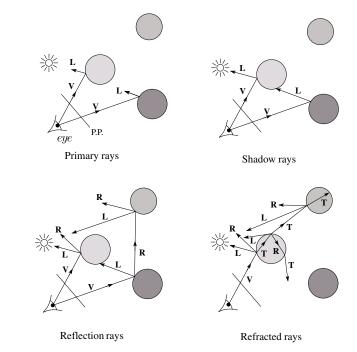


#### Algorithm:

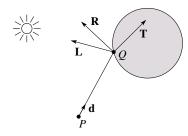
- 1. For each pixel, trace a  $\boldsymbol{primary\,ray}$  in direction  $\boldsymbol{V}$  to the first visible surface.
- 2. For each intersection, trace **secondary rays**:
  - Shadow rays in directions L; to light sources
  - Reflected ray in direction R.
  - Refracted ray or transmitted ray in direction T.

### Whitted algorithm (cont'd)

Let's look at this in stages:



## **Shading**



A ray is defined by an origin **P** and a unit direction **d** and is parameterized by *t*:

$$P + t\mathbf{d}$$

Let  $I(P, \mathbf{d})$  be the intensity seen along that ray. Then:

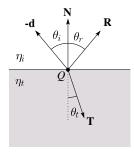
$$I(P, \mathbf{d}) = I_{\text{direct}} + I_{\text{reflected}} + I_{\text{transmitted}}$$

where

- I<sub>direct</sub> is computed from the Phong model
- $I_{\text{reflected}} = k_r I(Q, \mathbf{R})$
- $I_{\text{transmitted}} = k_t I(Q, \mathbf{T})$

Typically, we set  $k_r = k_s$  and  $k_t = 1 - k_s$ .

### **Reflection and transmission**



Law of reflection:

$$\theta_i = \theta_r$$

Snell's law of refraction:

$$\eta_{\rm i} \sin \theta_{\rm i} = \eta_{\rm t} \sin \theta_{\rm t}$$

where  $\eta_i$ ,  $\eta_t$  are indices of refraction.

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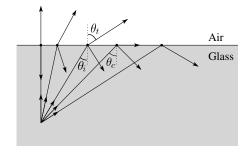
## **Total Internal Reflection**

The equation for the angle of refraction can be computed from Snell's law:

What happens when  $\eta_i > \eta_t$ ?

When  $\theta_t$  is exactly 90°, we say that  $\theta_t$  has achieved the "critical angle"  $\theta_c$ .

For  $\theta_l > \theta_c$ , no rays are transmitted, and only reflection occurs, a phenomenon known as "total internal reflection" or TIR.



#### **Error in Watt!!**

In order to compute the refracted direction, it is useful to compute the cosine of the angle of refraction in terms of the incident angle and the ratio of the indices of refraction.

On page 24 of Watt, he develops a formula for computing this cosine. Notationally, he uses  $\mu$  instead of  $\eta$  for the index of refraction in the text, but uses  $\eta$  in Figure 1.16(!?), and the angle of incidence is  $\phi$  and the angle of refraction is  $\theta$ .

Unfortunately, he makes a grave error in computing  $\cos \theta$ .

The last equation on page 24 should read:

$$\cos\theta = \sqrt{1 - \mu^2 (1 - \cos^2\phi)}$$

## Ray-tracing pseudocode

We build a ray traced image by casting rays through each of the pixels.

```
function tracelmage (scene):

for each pixel (i,j) in image

S = pixelToWorld(i,j)

P = COP

\mathbf{d} = (S - P)/||S - P||

I(i,j) = traceRay(scene, P, \mathbf{d})

end for
```

end function

## Ray-tracing pseudocode, cont'd

```
function traceRay(scene, P, d):
     (t, \mathbf{N}, mtrl) \leftarrow scene.intersect(P, \mathbf{d})
     Q \leftarrow \text{ray}(P, \mathbf{d}) \text{ evaluated at t}
     I = shade(\mathbf{q}, \mathbf{N}, mtrl, scene)
     \mathbf{R} = reflectDirection(\mathbf{N}, -\mathbf{d})
     I \leftarrow I + mtrl.k_r * traceRay(scene, Q, \mathbf{R})
     if ray is entering object then
           n_i = index_of_air
           n t = mtrl.index
     else
           n_i = mtrl.index
           n_t = index_of_air
     if (mtrl.k_t > 0 and notTIR(n_i, n_t, N, -d)) then
           T = refractDirection (n_i, n_t, N, -d)
           I \leftarrow I + mtrl.k_t * traceRay(scene, Q, T)
     end if
     return |
end function
```

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# **Terminating recursion**

**Q**: How do you bottom out of recursive ray tracing?

Possibilities:

## **Shading pseudocode**

Next, we need to calculate the color returned by the *shade* function.

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```
function shade (mtrl, scene, Q, N, d):
I \leftarrow mtrl.k_e + mtrl.k_a * scene->I_a
for each light source \ell do:
atten = \ell -> distanceAttenuation(Q) *
\ell -> shadowAttenuation(scene, Q)
I \leftarrow I + atten*(diffuse term + spec term)
end for
return \ I
end function
```

### **Shadow attenuation**

Computing a shadow can be as simple as checking to see if a ray makes it to the light source.

For a point light source:

**function** *PointLight*::shadowAttenuation(scene, P)

 $\mathbf{d} = (\ell.\text{position } - P).normalize()$ 

 $(t, \mathbf{N}, mtrl) \leftarrow scene.intersect(P, \mathbf{d})$ 

 $Q \leftarrow \text{ray}(t)$ 

**if** *Q* is before the light source **then**:

atten = 0

else

atten = 1

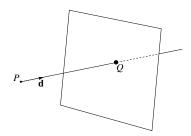
end if

return atten

end function

**Q**: What if there are transparent objects along a path to the light source?

# **Ray-plane intersection**



We can write the equation of a plane as:

$$ax + by + cz + d = 0$$

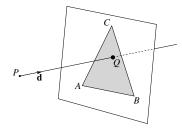
The coefficients a, b, and c form a vector that is normal to the plane,  $\mathbf{n} = [a \ b \ c]^{\mathsf{T}}$ . Thus, we can rewrite the plane equation as:

We can solve for the intersection parameter (and thus the point):

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## **Ray-triangle intersection**



To intersect with a triangle, we first solve for the equation of its supporting plane:

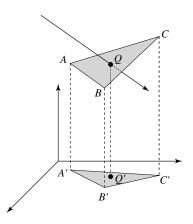
Then, we need to decide if the point is inside or outside of the triangle.

Solution 1: compute barycentric coordinates from 3D points.

What do you do with the barycentric coordinates?

## **Ray-triangle intersection**

Solution 2: project down a dimension and compute barycentric coordinates from 2D points.



Why is solution 2 possible? Why is it legal? Why is it desirable? Which axis should you "project away"?

### **Interpolating vertex properties**

The barycentric coordinates can also be used to interpolate vertex properties such as:

- material properties
- texture coordinates
- normals

For example:

$$k_d(Q) = \alpha k_d(A) + \beta k_d(B) + \gamma k_d(C)$$

Interpolating normals, known as Phong interpolation, gives triangle meshes a smooth shading appearance.

### **Epsilons**

Due to finite precision arithmetic, we do not always get the exact intersection at a surface.

**Q**: What kinds of problems might this cause?

**Q**: How might we resolve this?

## Intersecting with xformed geometry

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In general, objects will be placed using transformations. What if the object being intersected were transformed by a matrix M?

Apply M<sup>-1</sup> to the ray first and intersect in object (local) coordinates!

## Intersecting with xformed geometry

The intersected normal is in object (local) coordinates. How do we transform it to world coordinates?

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